Plans and possibilities for fielding HF receivers at HAARP

Brett Isham Interamerican University of Puerto Rico

Collaborators

Jim LaBelle and Matt Broughton Dartmouth College

HAARP/Resonance Workshop 8-9 November 2011 University of Maryland, College Park, Maryland, USA

Contributors (alphabetical order)

Vasyl Belyey, University of Tromsø, Norway Jan Bergman, Swedish Institute of Space Physics, Sweden Paul Bernhardt, Naval Research Laboratory, USA Richard Brittain, Dartmouth College, USA Matt Broughton, Dartmouth College, USA Terry Bullett, National Oceanic and Atmospheric Administration, USA Nick Bunch, Stanford University, USA Jorge Chau, Jicamarca Radio Observatory, Peru Lars Daldorff, University of Michigan, USA Paul Gallop, Soffari Ltd., Reading, UK Patrick Guio, University College London, London, UK Dave Hysell, Cornell University, USA Jim LaBelle, Dartmouth College, USA Robert Livingston, Scion Associates, USA David McGaw, Dartmouth College, USA Siavoush Mohammadi, UUSU, Sweden Walter Puccio, Swedish Institute of Space Physics, Sweden Mike Rietveld, EISCAT Scientific Association, Norway Evgeny Tereshchenko, Polar Geophysical Institute, Russia Bo Thidé, Swedish Institute of Space Physics, Sweden Mike Trimpi, Dartmouth College, USA Bill Wright, University of Colorado, USA Roman Yurik, Polar Geophysical Institute, Russia Robert Zimmerman, Interamerican University of Puerto Rico, USA

Digital radio receivers – multi-purpose

SEE

spectra polarization direction angle imaging

Radio phase modes – photon orbital angular momentum (OAM)

Pump wave – verify transmission (time, frequency)

Radar receiver – HF, ionosonde

Satellite beacon scintillation receiver – VHF/UHF

Natural radio emissions – unclamped pump, wide bandwidth

Digital radio receivers

Wide band

high-speed data transfer

Multi-channel

four or eight channels

Coherent

coherent operation at multiple sites

Low-maintenance

easily configurable, unattended operation, remotely controllable modular, identical swapable parts

Radio emission measurements

Single site

stimulated radio emissions (SEE) power (time, frequency) polarization (time, frequency)

Single site – coherent receivers

direction angle

Multiple sites

aspect angle dependence

Multiple sites – coherent receivers

imaging radio phase modes (photon orbital angular momentum)

Collaborations for receiver development

Swedish Institute of Space Physics (Bo Thidé et al., 2006-present)

12 receivers

4 channels per receiver

16-bit ADC, 125-MHz sampling

control and data via Ethernet (two 1-Gbps lines)

coherence via GPS with advanced processing (Thidé et al.)

produced by BitSim AB, Stockholm, Sweden

Scion Associates (Bob Livingston et al., 2010-present)

8 receivers

8 channels per receiver

16-bit ADC, 100-MHz sampling

control via USB-2, data via USB-3 (5 Gbps) or Thunderbolt (2x10 Gbps) coherence via GPS-disciplined rubidium (frequency accuracy $\pm 1 \times 10^{-8}$) produced by Scion Associates, Port Townsend, Washington, USA



Schematic of SEE spectra

Schematic diagram of SEE spectra for variations in HF pump frequency relative to the gyroharmonic frequencies (n > 3).

Note the asymmetries in the individual spectra as well as in spectra above and below the gyroharmonics.

Adapted from Leyser (2001)

SEE spectra observed at Sura



Frolov et al. (2001)



Fig. 4. Stimulated electromagnetic emission spectra for a 5.6 MHz pump wave pointed toward the magnetic zenith from HAARP. A down shifted SBS-2 line may be masked by the low-frequency side of the strong SBS-1 line.

Comparison with theory: Results and predictions

New simulation codes include EM emission

fluid, kinetic, particle, hybrid, ...

1-D, 2-D, and 3-D

Few if any citable predictions, e.g.:

Scales et al. (1997): qualitative similarity with observations Mjølhus (1998): consistent with observations Xi and Scales (2001): exhibit many properties of observations Eliasson and Stenflo (2010): spectra agree with observations

SEE experiments lead theory...



Radio observations with new Arecibo HF



Radio aspect angle array for Arecibo HF experiments

> Arecibo Bayamón Culebra Saint Croix

0 km, 90° • 50 km, 74° • 160 km, 48° • 220 km, 39° • 580 km, 17° •

Polarization: multi-channel receivers and electric triad antennas

Thidé, Bergman, Isham, et al.

Magnetic triad antenna

Thidé, Bergman, Isham, et al.





Radio direction-finding



Dartmouth College direction-finding array

Toolik Lake, Alaska August 2006

(photo by Nick Bunch)

Direction angle of stimulated radio emissions (SEE)

Broad upshifted maximum (BUM1)

0 to +100 kHz from 5.423-MHz pump (~ 4 *f*ce) 2 min on, 2 min off

HF: 14.0°S, 17.5°S BUM1: 16°S, 20°S

Ramfjordmoen 5 October 2004 10:55 to 11:05 UT



Location of downshifted peak (DP) for pump beam at 0°, 7°, 14°, and 21° S



Tereshchenko et al. (2006)



buildings +-200 m EW

Simulated 1-D radio image



Belyey and Isham

Photon orbital angular momentum (radio phase modes) at HAARP



Radio Phase Modes





Summary

New digital receivers

Aspect angle

Polarization

Direction finding

Imaging

Radio phase modes

Start simply, add complexity

Outside of HF campaigns, extend Dartmouth network

Collaborations in the use of the receivers are welcome!

Contact Brett Isham at <<u>brettisham@gmail.com</u>>